

Lect 2. Boltzmann and Maxwell

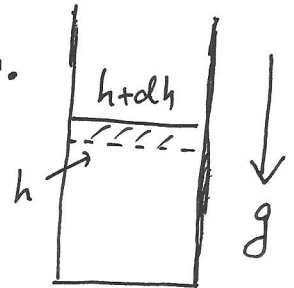
distributions

1. Boltzmann distribution — gas in a gravity field
2. From Boltzmann \longrightarrow Maxwell distribution
3. Gaussian integral
4. Probability distribution function

§ Boltzmann distribution

- Consider a gas in the gravity field at thermal equilibrium.

The temperature is the same, but the density may not be uniform. The pressure difference between the height



h and h+dh, is due to the gravity of gas molecule within the thickness of dh.

$$p(h) - p(h+dh) = mng dh = - \frac{dp}{dh} dh$$

$$p(h) = n(h) k_B T \Rightarrow \frac{dp}{dh} = \frac{dn}{dh} k_B T \Rightarrow \frac{dn}{dh} = - \frac{1}{k_B T} mng$$

$$\frac{1}{n} \frac{dn}{dh} = - \frac{mg}{k_B T} \Rightarrow \frac{d \ln n}{dh} = - \frac{mg}{k_B T}$$

$$n = n_0 e^{-\frac{mgh}{k_B T}}$$

Here, "mgh" can be viewed as the gravity potential energy, which can be generalized to other potential, say, electric potential energy, etc.

Then a characteristic height $h_0 = \frac{k_B T}{mg}$ can be defined.

$$h_0 = \frac{k_B T}{mc^2} \cdot \frac{c^2}{g}$$

$$T = 300 \text{ K}$$

$$m: \text{O}_2 \approx 32 \text{ GeV} \approx 3 \times 10^{10} \text{ eV} \sim 3 \times 10^{14} \text{ K}$$

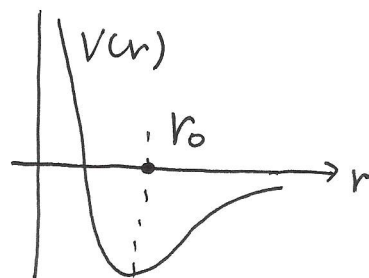
$$\Rightarrow h_0 \approx \frac{3 \times 10^2}{3 \times 10^{14}} \times 10^{16}$$

$$c^2/g \approx \frac{9 \times 10^{16}}{10} \sim 10^{16} \text{ m}$$

$\sim 10^4 \text{ m}$. \Rightarrow O_2 density reduces to $1/e \approx 37\%$ at 10 km height!

- intermolecular interaction

In gases, the interactions among gas molecules can often be neglected, but in liquids they cannot be neglected. The interactions are typically pair-wised,



Then the probability for find a configuration with the intermolecule distance r_{ij} ,

$$P(\{r_{ij}\}) \propto e^{-\sum_{i,j} V(r_{ij})/k_B T}$$

If $k_B T \ll |V(r_0)|$, the relative probability to find a molecules at small distances around r_0 will be large. The system could be in a

liquid state. On the other hand, if $k_B T \geq |V(r_0)|$, the molecules will tend to go at large distance, since there are more volume,

and $\int d^3\vec{r}_1 \dots d^3\vec{r}_N$ (the integral measures are large at r_{ij} is large).

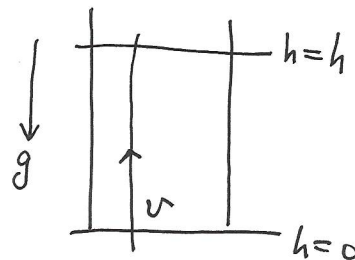
Say, for 2 molecules $\int d^3\vec{r}_1 \int d^3\vec{r}_2 = \int d^3\vec{R} \int d^3\vec{r}$, and $\int d^3\vec{r} = \int r^2 dr \int d\Omega$.

Hence, as varying temperature, the interaction $\sum_{i,j} V(r_{ij})$ could have different effect. At high temperatures, the inter-molecular distance is large, the system is a gas. While as temperature goes down, interactions will reorganize them to favor $V(r_{ij})$ reaching minima. The system would go to liquid, and Even solid. This is a complicated problem - "many body problem" or "Condensed matter physics".

§ Maxwell distribution

We have derived Maxwell's distribution in HW. Now let us examine it from a different viewpoint. Consider a gas in the gravity field.

- Define $\mathcal{L}(h, v_z)$ is the number of molecules ^{per unit time} per unit volume passing the horizontal plane at ^{the} height of h with velocity larger than v_z .
_{the vertical}

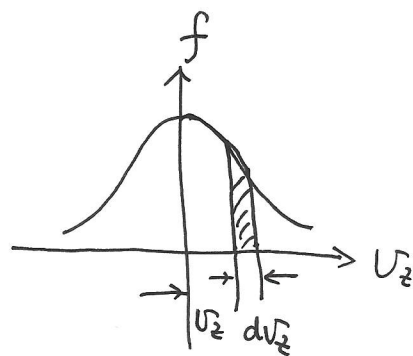


- Define the distribution function $f(v_z)$, satisfying

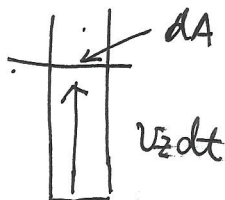
$$\int_{-\infty}^{+\infty} dv_z f(v_z) = 1,$$

$f(v_z) dv_z$: the probability for a molecule's v_z lying from v_z to $v_z + dv_z$

$$d(\# \text{ of molecules}) = dA dt v_z n(h) f(v_z) dv_z$$



$$\mathcal{L}(h, v_z) = n(h) \int_{v_z}^{+\infty} dv_z v_z f(v_z)$$



Hence,

$$\frac{\mathcal{L}(h, v_z=0)}{\mathcal{L}(h=0, v_z=0)} = \frac{n(h)}{n(h=0)}$$

where $n(h)$ is the molecule density at the height h .

At $h=0$, if the molecule's velocity $v_z > u = \sqrt{2gh}$, then it can pass the height h with non-zero vertical velocity > 0 . Then

$$l(0, u) = l(h, 0) \quad \text{with } \frac{1}{2}mu^2 = mgh \text{ and } u > 0.$$

$$\Rightarrow \frac{l(0, u)}{l(0, 0)} = \frac{l(h, 0)}{l(0, 0)} = \frac{n(h)}{n(0)} = e^{-mgh/k_B T} = e^{-\frac{1}{2}mu^2/k_B T}$$

$$\Rightarrow \text{define } F(u) = \int_u^{+\infty} dv_z v_z f(v_z) = C e^{-\frac{1}{2}mu^2/k_B T}$$

with $C = F(u=0)$ is a constant.

$$\text{Then } \frac{dF(u)}{du} = -u f(u) = -mu C e^{-\frac{1}{2}mu^2/k_B T}$$

$$\Rightarrow f(u) \propto e^{-\frac{1}{2}mu^2/k_B T}$$

$$\text{or } f(v_z) = \frac{e^{-\frac{\beta}{2}mv_z^2}}{\int_{-\infty}^{+\infty} dv_z e^{-\frac{\beta}{2}mv_z^2}} \quad \text{with } \beta = \frac{1}{k_B T}$$

* Knowledge of Gaussian integral

$$\begin{aligned} a &= \int_{-\infty}^{+\infty} dx e^{-x^2} \Rightarrow a^2 = \int_{-\infty}^{+\infty} dx \int_{-\infty}^{+\infty} dy e^{-(x^2+y^2)} = \int_0^{+\infty} r dr \int_0^{2\pi} d\theta e^{-r^2} \\ &\Rightarrow \int_{-\infty}^{+\infty} dx e^{-x^2} = \sqrt{\pi} \end{aligned}$$

$$\int_{-\infty}^{+\infty} dx e^{-ax^2} = \int_{-\infty}^{+\infty} \frac{dy}{\sqrt{a}} e^{-y^2} = \sqrt{\frac{\pi}{a}} \Rightarrow \int_{-\infty}^{+\infty} dv_z e^{-\frac{\beta m}{2} v_z^2} = \sqrt{\frac{2\pi}{m\beta}}$$

$$y = \sqrt{a}x$$

define $f(a) = \int_{-\infty}^{+\infty} dx e^{-ax^2} \Rightarrow \frac{df(a)}{da} = - \int_{-\infty}^{+\infty} dx x^2 e^{-ax^2} = -\frac{1}{2} \frac{\sqrt{\pi}}{a^{3/2}}$

$$\Rightarrow \int_{-\infty}^{+\infty} dx x^2 e^{-ax^2} = \frac{\pi^{1/2}}{2a^{3/2}}$$

* Maxwell distribution summary

$$f(v_z) = \left(\frac{2\pi}{m\beta}\right)^{-1/2} e^{-\frac{\beta}{2} m v_z^2}$$


$$\Rightarrow f(v_x, v_y, v_z) = \left(\frac{m}{2\pi k_B T}\right)^{3/2} e^{-\frac{\beta}{2} m (v_x^2 + v_y^2 + v_z^2)}$$

$$\int dv_x = \int \frac{dp_x}{m}$$

or we switch to momentum p_x, p_y, p_z

$$f(p_x, p_y, p_z) = \left(\frac{1}{2\pi m k_B T}\right)^{3/2} e^{-\frac{\beta}{2m} (p_x^2 + p_y^2 + p_z^2)}$$

$$\int_{-\infty}^{+\infty} dp_x dp_y dp_z f(p_x, p_y, p_z) = 1$$

$f(p_x, p_y, p_z) dp_x dp_y dp_z$ 
 probability lying in this volume element
 in momentum space

④ Distribution of speed u

$$\left(\frac{m}{2\pi k_B T}\right)^{3/2} \int dv_x dv_y dv_z e^{-\frac{m}{2k_B T}(v_x^2 + v_y^2 + v_z^2)} = 1$$

$$\left(\frac{m}{2\pi k_B T}\right)^{3/2} \int_0^{\infty} v^2 dv \int d\Omega e^{-\frac{m}{2k_B T} v^2} = \int_0^{\infty} dv e^{-\frac{m v^2}{2k_B T}} \cdot v^2 \cdot (4\pi) \left(\frac{m}{2\pi k_B T}\right)^{3/2} = 1$$

$$\Rightarrow f(v) = v^2 e^{-\frac{m v^2}{2k_B T}} \cdot \sqrt{\frac{2}{\pi}} \left(\frac{m}{k_B T}\right)^{3/2}$$