

Lecture 11.5 Landau Fermi liquid (V)

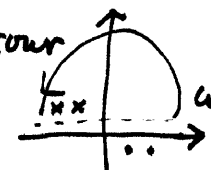
§ More on Luttinger theorem — Luttinger & Ward

Another expression:
$$\frac{N}{V} = 2i \int \left(\frac{\partial}{\partial \omega} \ln G(p) - G(p) \frac{\partial}{\partial \omega} \Sigma(p) \right) e^{i\omega\tau} \frac{d^4p}{(2\pi)^4}$$

Proof:

$$\begin{aligned} \frac{\partial}{\partial \omega} \ln G(p) - G(p) \frac{\partial}{\partial \omega} \Sigma(p) &= G^{-1}(p) \frac{\partial}{\partial \omega} G(p) - G \frac{\partial}{\partial \omega} \Sigma = (\omega - \xi - \Sigma) \frac{\partial}{\partial \omega} G \\ - G \frac{\partial}{\partial \omega} \Sigma &= \omega \frac{\partial}{\partial \omega} G - \xi \frac{\partial}{\partial \omega} G - \frac{\partial}{\partial \omega} (\Sigma G) = \frac{\partial}{\partial \omega} (\omega G) - \frac{\partial}{\partial \omega} (\Sigma G) \\ &\quad - \frac{\partial}{\partial \omega} (\xi G) - G \\ &= \frac{\partial}{\partial \omega} [(\omega - \xi - \Sigma) G] - G \Rightarrow \boxed{\frac{N}{V} = -2i \int G e^{i\omega\tau} \frac{d^4p}{(2\pi)^4}} \end{aligned}$$

The second term of the integral can be proved to be zero. Let us postpone

its proof for a while.
$$\frac{N}{V} = 2i \int \frac{\partial}{\partial \omega} \ln G(p) e^{i\omega\tau} \frac{d^4p}{(2\pi)^4}$$
 

G is not analytic. $G_R = \begin{cases} G(\epsilon) & \epsilon > 0 \\ G^*(\epsilon) & \epsilon < 0 \end{cases}$, G_R is analytic for ω in the upper plane

$$\begin{aligned} \text{So } \frac{N}{V} &= 2i \left[\int_0^\infty \frac{d\omega}{2\pi} \int \frac{d^3p}{(2\pi)^3} \frac{\partial}{\partial \omega} \ln G_R(p) + \int_{-\infty}^0 \frac{d\omega}{2\pi} \int \frac{d^3p}{(2\pi)^3} \frac{\partial}{\partial \omega} \ln G(p) \right] \\ &= 2i \left[\int_{-\infty}^{+\infty} \frac{d\omega}{2\pi} \int \frac{d^3p}{(2\pi)^3} \frac{\partial}{\partial \omega} \ln G_R(p) + \int_{-\infty}^0 \frac{d\omega}{2\pi} \int \frac{d^3p}{(2\pi)^3} \frac{\partial}{\partial \omega} \ln \frac{G(p, \omega)}{G^*(p, \omega)} \right] \\ &\quad \parallel \\ &\quad 0 \\ &= \frac{2i}{2\pi} \int \frac{d^3p}{(2\pi)^3} \ln \frac{G(p, \omega)}{G^*(p, \omega)} \Big|_{-\infty}^0 = -\frac{2}{\pi} \int \frac{d^3p}{(2\pi)^3} [\varphi(\omega) - \varphi(-\infty)], \end{aligned}$$

where φ denotes the phase of G . We need to consider the variation of φ from $\omega = -\infty$ to $\omega = 0$. From the Lehmann Rep. we have

$$G(\omega) = e^{\beta\mu} \sum_{n,m} \langle n | \psi | m \rangle \langle m | \psi^\dagger | n \rangle \left\{ \frac{e^{-\beta E_n}}{\omega + E_n - E_m + i\eta} + \frac{e^{-\beta E_m}}{\omega + E_n - E_m - i\eta} \right\}$$

$$\xrightarrow{T=0K} = \sum_m \frac{\langle 0 | \psi | m \rangle \langle m | \psi^\dagger | 0 \rangle}{\omega - E_m + i\eta} + \sum_n \frac{\langle 0 | \psi^\dagger | n \rangle \langle n | \psi | 0 \rangle}{\omega + E_n - i\eta}$$

$$= \int_0^{+\infty} dE \frac{A(E)}{\omega - E + i\eta} + \frac{B(E)}{\omega + E - i\eta},$$

A, B are positive

where $A(E) = \sum_m \langle 0 | \psi | m \rangle \langle m | \psi^\dagger | 0 \rangle \delta(E - E_m)$

$$B(E) = \sum_n \langle 0 | \psi^\dagger | n \rangle \langle n | \psi | 0 \rangle \delta(E - E_n).$$

As $\omega \rightarrow \infty \Rightarrow G(\omega) \rightarrow \frac{1}{\omega} \int_0^{+\infty} (A(E) + B(E)) dE$

$$= \sum_{m,n} \frac{1}{\omega} \{ \langle 0 | \psi | m \rangle \langle m | \psi^\dagger | 0 \rangle + \langle 0 | \psi^\dagger | n \rangle \langle n | \psi | 0 \rangle \} = \frac{1}{\omega} \langle 0 | \psi \psi^\dagger + \psi^\dagger \psi | 0 \rangle$$

$$= \frac{1}{\omega}$$

i.e.

$G(\omega) \rightarrow \frac{1}{\omega}, \text{ as } \omega \rightarrow \infty$

$$\text{Re } G(\omega) = \mathcal{P} \int_0^{+\infty} dE \frac{A(E)}{\omega - E} + \frac{B(E)}{\omega + E}$$

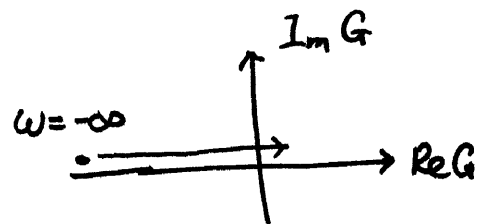
$$\text{Im } G(\omega) = \begin{cases} -\pi A(\omega) & \text{for } \omega > 0 \\ \pi B(-\omega) & \text{for } \omega < 0 \end{cases}.$$

As $\omega \rightarrow -\infty$, $\text{Re } G(\omega) \rightarrow -\frac{1}{\omega} < 0$. $\text{Im } G(\omega)$ at $\omega < 0$, describe the hole excitation below the Fermi surface. If we assume that

the hole-excitation is bounded from below, then $\text{Im} G(\omega) > 0$ ③

decays faster, $\Rightarrow \varphi(\omega \rightarrow -\infty) = \pi$.

↳ towards 0



We set $\text{Im} G(\omega) = 0$ at $\omega = 0$, because all the ϵ_n, E_m in the Lehmann Rep > 0 . The φ of $G(\omega=0)$ is determined by the real part $\text{Re} G(\omega=0)$. If $\text{Re} G(\omega=0) < 0 \Rightarrow \varphi = \pi$,
 $\text{Re} G(\omega=0) > 0 \Rightarrow \varphi = 0$.

$$\Rightarrow \frac{N}{V} = 2 \int \frac{d^3 p}{(2\pi)^3} \left[\text{Re} G(p, \omega=0) > 0 \right]$$

← The region $G(p, \omega=0)$ is bound by a surface of either zero or divergence.

Vanishing of $G(p, \omega=0)$ corresponds to $\Sigma \rightarrow \infty$, which

corresponds to superconductivity $G = \frac{u_k^2}{\omega - \sqrt{E_k^2 + \Delta^2}} + \frac{v_k^2}{\omega + \sqrt{E_k^2 + \Delta^2}}$.

$\Rightarrow G(0, k) = \frac{-u_k^2 + v_k^2}{E_k}$, (we do not consider this possibility here).

On the other hand, $G(p, \omega=0) = \frac{z}{\omega - \xi_p + i\eta} + \dots \rightarrow +\infty$

corresponds to location of $\xi_p = 0$, i.e. the location of FS.

$G(p, \omega=0) \rightarrow -\frac{z}{\xi_p} \rightarrow +\infty \quad \xi_p < 0$
 $\rightarrow -\infty \quad \text{for } \xi_p > 0$.

This proof does not assume the isotropy of FS. Interaction may deform the shape but not its volume!